

# Evolution of $\pi^0$ suppression in Au+Au collisions from $\sqrt{s_{NN}} = 39$ to 200 GeV

A. Adare,<sup>11</sup> S. Afanasiev,<sup>26</sup> C. Aidala,<sup>34</sup> N.N. Ajitanand,<sup>54</sup> Y. Akiba,<sup>48,49</sup> R. Akimoto,<sup>10</sup> H. Al-Ta'ani,<sup>43</sup> J. Alexander,<sup>54</sup> A. Angerami,<sup>12</sup> K. Aoki,<sup>48</sup> N. Apadula,<sup>55</sup> Y. Aramaki,<sup>10,48</sup> H. Asano,<sup>31,48</sup> E.C. Aschenauer,<sup>6</sup> E.T. Atomssa,<sup>55</sup> T.C. Awes,<sup>45</sup> B. Azmoun,<sup>6</sup> V. Babintsev,<sup>22</sup> M. Bai,<sup>5</sup> B. Bannier,<sup>55</sup> K.N. Barish,<sup>7</sup> B. Bassalleck,<sup>42</sup> S. Bathe,<sup>4,49</sup> V. Baublis,<sup>47</sup> S. Baumgart,<sup>48</sup> A. Bazilevsky,<sup>6</sup> R. Belmont,<sup>59</sup> A. Berdnikov,<sup>51</sup> Y. Berdnikov,<sup>51</sup> X. Bing,<sup>44</sup> D.S. Blau,<sup>30</sup> K. Boyle,<sup>49</sup> M.L. Brooks,<sup>34</sup> H. Buesching,<sup>6</sup> V. Bumazhnov,<sup>22</sup> S. Butsyk,<sup>42</sup> S. Campbell,<sup>55</sup> P. Castera,<sup>55</sup> C.-H. Chen,<sup>55</sup> C.Y. Chi,<sup>12</sup> M. Chiu,<sup>6</sup> I.J. Choi,<sup>23</sup> J.B. Choi,<sup>9</sup> S. Choi,<sup>53</sup> R.K. Choudhury,<sup>3</sup> P. Christiansen,<sup>36</sup> T. Chujo,<sup>58</sup> O. Chvala,<sup>7</sup> V. Cianciolo,<sup>45</sup> Z. Citron,<sup>55</sup> B.A. Cole,<sup>12</sup> M. Connors,<sup>55</sup> M. Csanád,<sup>16</sup> T. Csörgő,<sup>61</sup> S. Dairaku,<sup>31,48</sup> A. Datta,<sup>38</sup> M.S. Daugherty,<sup>1</sup> G. David,<sup>6</sup> A. Denisov,<sup>22</sup> A. Deshpande,<sup>49,55</sup> E.J. Desmond,<sup>6</sup> K.V. Dharmawardane,<sup>43</sup> O. Dietzsch,<sup>52</sup> L. Ding,<sup>25</sup> A. Dion,<sup>25</sup> M. Donadelli,<sup>52</sup> O. Drapier,<sup>32</sup> A. Drees,<sup>55</sup> K.A. Drees,<sup>5</sup> J.M. Durham,<sup>55</sup> A. Durum,<sup>22</sup> L. D'Orazio,<sup>37</sup> S. Edwards,<sup>5</sup> Y.V. Efremenko,<sup>45</sup> T. Engelmores,<sup>12</sup> A. Enokizono,<sup>45</sup> S. Esumi,<sup>58</sup> K.O. Eyser,<sup>7</sup> B. Fadern,<sup>39</sup> D.E. Fields,<sup>42</sup> M. Finger,<sup>8</sup> M. Finger, Jr.,<sup>8</sup> F. Fleuret,<sup>32</sup> S.L. Fokin,<sup>30</sup> J.E. Frantz,<sup>44</sup> A. Franz,<sup>6</sup> A.D. Frawley,<sup>18</sup> Y. Fukao,<sup>48</sup> T. Fusayasu,<sup>41</sup> K. Gainey,<sup>1</sup> C. Gal,<sup>55</sup> A. Garishvili,<sup>56</sup> I. Garishvili,<sup>33</sup> A. Glenn,<sup>33</sup> X. Gong,<sup>54</sup> M. Gonin,<sup>32</sup> Y. Goto,<sup>48,49</sup> R. Granier de Cassagnac,<sup>32</sup> N. Grau,<sup>12</sup> S.V. Greene,<sup>59</sup> M. Grosse Perdekamp,<sup>23</sup> T. Gunji,<sup>10</sup> L. Guo,<sup>34</sup> H.-Å. Gustafsson,<sup>36,\*</sup> T. Hachiya,<sup>48</sup> J.S. Haggerty,<sup>6</sup> K.I. Hahn,<sup>17</sup> H. Hamagaki,<sup>10</sup> J. Hanks,<sup>12</sup> K. Hashimoto,<sup>48,50</sup> E. Haslum,<sup>36</sup> R. Hayano,<sup>10</sup> X. He,<sup>19</sup> T.K. Hemmick,<sup>55</sup> T. Hester,<sup>7</sup> J.C. Hill,<sup>25</sup> R.S. Hollis,<sup>7</sup> K. Homma,<sup>21</sup> B. Hong,<sup>29</sup> T. Horaguchi,<sup>58</sup> Y. Hori,<sup>10</sup> S. Huang,<sup>59</sup> T. Ichihara,<sup>48,49</sup> H. Iinuma,<sup>28</sup> Y. Ikeda,<sup>48,58</sup> J. Imrek,<sup>15</sup> M. Inaba,<sup>58</sup> A. Iordanova,<sup>7</sup> D. Isenhowe,<sup>1</sup> M. Issah,<sup>59</sup> A. Isupov,<sup>26</sup> D. Ivanischev,<sup>47</sup> B.V. Jacak,<sup>55,†</sup> M. Javani,<sup>19</sup> J. Jia,<sup>6,54</sup> X. Jiang,<sup>34</sup> B.M. Johnson,<sup>6</sup> K.S. Joo,<sup>40</sup> D. Jouan,<sup>46</sup> J. Kamin,<sup>55</sup> S. Kaneti,<sup>55</sup> B.H. Kang,<sup>20</sup> J.H. Kang,<sup>62</sup> J.S. Kang,<sup>20</sup> J. Kapustinsky,<sup>34</sup> K. Karatsu,<sup>31,48</sup> M. Kasai,<sup>48,50</sup> D. Kawall,<sup>38,49</sup> A.V. Kazantsev,<sup>30</sup> T. Kempel,<sup>25</sup> A. Khanzadeev,<sup>47</sup> K.M. Kijima,<sup>21</sup> B.I. Kim,<sup>29</sup> C. Kim,<sup>29</sup> D.J. Kim,<sup>27</sup> E.J. Kim,<sup>9</sup> H.J. Kim,<sup>62</sup> K.-B. Kim,<sup>63</sup> Y.-J. Kim,<sup>23</sup> Y.K. Kim,<sup>20</sup> E. Kinney,<sup>11</sup> Á. Kiss,<sup>16</sup> E. Kistenev,<sup>6</sup> J. Klatsky,<sup>18</sup> D. Kleinjan,<sup>7</sup> P. Kline,<sup>55</sup> Y. Komatsu,<sup>10</sup> B. Komkov,<sup>47</sup> J. Koster,<sup>23</sup> D. Kotchetkov,<sup>44</sup> D. Kotov,<sup>51</sup> A. Král,<sup>13</sup> F. Krizek,<sup>27</sup> G.J. Kunde,<sup>34</sup> K. Kurita,<sup>48,50</sup> M. Kurosawa,<sup>48</sup> Y. Kwon,<sup>62</sup> G.S. Kyle,<sup>43</sup> R. Lacey,<sup>54</sup> Y.S. Lai,<sup>12</sup> J.G. Lajoie,<sup>25</sup> A. Lebedev,<sup>25</sup> B. Lee,<sup>20</sup> D.M. Lee,<sup>34</sup> J. Lee,<sup>17</sup> K.B. Lee,<sup>29</sup> K.S. Lee,<sup>29</sup> S.H. Lee,<sup>55</sup> S.R. Lee,<sup>9</sup> M.J. Leitch,<sup>34</sup> M.A.L. Leite,<sup>52</sup> M. Leitgab,<sup>23</sup> B. Lewis,<sup>55</sup> S.H. Lim,<sup>62</sup> L.A. Linden Levy,<sup>11</sup> A. Litvinenko,<sup>26</sup> M.X. Liu,<sup>34</sup> B. Love,<sup>59</sup> C.F. Maguire,<sup>59</sup> Y.I. Makdisi,<sup>5</sup> M. Makek,<sup>60</sup> A. Malakhov,<sup>26</sup> A. Manion,<sup>55</sup> V.I. Manko,<sup>30</sup> E. Mannel,<sup>12</sup> S. Masumoto,<sup>10</sup> M. McCumber,<sup>11</sup> P.L. McGaughey,<sup>34</sup> D. McGlinchey,<sup>18</sup> C. McKinney,<sup>23</sup> M. Mendoza,<sup>7</sup> B. Meredith,<sup>23</sup> Y. Miake,<sup>58</sup> T. Mibe,<sup>28</sup> A.C. Mignerey,<sup>37</sup> A. Milov,<sup>60</sup> D.K. Mishra,<sup>3</sup> J.T. Mitchell,<sup>6</sup> Y. Miyachi,<sup>48,57</sup> S. Miyasaka,<sup>48,57</sup> A.K. Mohanty,<sup>3</sup> H.J. Moon,<sup>40</sup> D.P. Morrison,<sup>6</sup> S. Motschwiller,<sup>39</sup> T.V. Moukhanova,<sup>30</sup> T. Murakami,<sup>31,48</sup> J. Murata,<sup>48,50</sup> T. Nagae,<sup>31</sup> S. Nagamiya,<sup>28</sup> J.L. Nagle,<sup>11</sup> M.I. Nagy,<sup>61</sup> I. Nakagawa,<sup>48,49</sup> Y. Nakamiya,<sup>21</sup> K.R. Nakamura,<sup>31,48</sup> T. Nakamura,<sup>48</sup> K. Nakano,<sup>48,57</sup> C. Nattrass,<sup>56</sup> A. Nederlof,<sup>39</sup> M. Nihashi,<sup>21,48</sup> R. Nouicer,<sup>6,49</sup> N. Novitzky,<sup>27</sup> A.S. Nyanin,<sup>30</sup> E. O'Brien,<sup>6</sup> C.A. Ogilvie,<sup>25</sup> K. Okada,<sup>49</sup> A. Oskarsson,<sup>36</sup> M. Ouchida,<sup>21,48</sup> K. Ozawa,<sup>10</sup> R. Pak,<sup>6</sup> V. Papavassiliou,<sup>43</sup> B.H. Park,<sup>20</sup> I.H. Park,<sup>17</sup> S.K. Park,<sup>29</sup> S.F. Pate,<sup>43</sup> L. Patel,<sup>19</sup> H. Pei,<sup>25</sup> J.-C. Peng,<sup>23</sup> H. Pereira,<sup>14</sup> V. Peresedov,<sup>26</sup> D.Yu. Peressounko,<sup>30</sup> R. Petti,<sup>55</sup> C. Pinkenburg,<sup>6</sup> R.P. Pisani,<sup>6</sup> M. Proissl,<sup>55</sup> M.L. Purschke,<sup>6</sup> H. Qu,<sup>1</sup> J. Rak,<sup>27</sup> I. Ravinovich,<sup>60</sup> K.F. Read,<sup>45,56</sup> R. Reynolds,<sup>54</sup> V. Riabov,<sup>47</sup> Y. Riabov,<sup>47</sup> E. Richardson,<sup>37</sup> D. Roach,<sup>59</sup> G. Roche,<sup>35</sup> S.D. Rolnick,<sup>7</sup> M. Rosati,<sup>25</sup> P. Rukoyatkin,<sup>26</sup> B. Sahlmueller,<sup>55</sup> N. Saito,<sup>28</sup> T. Sakaguchi,<sup>6</sup> V. Samsonov,<sup>47</sup> M. Sano,<sup>58</sup> M. Sarsour,<sup>19</sup> S. Sawada,<sup>28</sup> K. Sedgwick,<sup>7</sup> R. Seidl,<sup>48,49</sup> A. Sen,<sup>19</sup> R. Seto,<sup>7</sup> D. Sharma,<sup>60</sup> I. Shein,<sup>22</sup> T.-A. Shibata,<sup>48,57</sup> K. Shigaki,<sup>21</sup> M. Shimomura,<sup>58</sup> K. Shoji,<sup>31,48</sup> P. Shukla,<sup>3</sup> A. Sickles,<sup>6</sup> C.L. Silva,<sup>25</sup> D. Silvermyr,<sup>45</sup> K.S. Sim,<sup>29</sup> B.K. Singh,<sup>2</sup> C.P. Singh,<sup>2</sup> V. Singh,<sup>2</sup> M. Slunečka,<sup>8</sup> R.A. Soltz,<sup>33</sup> W.E. Sondheim,<sup>34</sup> S.P. Sorensen,<sup>56</sup> M. Soumya,<sup>54</sup> I.V. Sourikova,<sup>6</sup> P.W. Stankus,<sup>45</sup> E. Stenlund,<sup>36</sup> M. Stepanov,<sup>38</sup> A. Ster,<sup>61</sup> S.P. Stoll,<sup>6</sup> T. Sugitate,<sup>21</sup> A. Sukhanov,<sup>6</sup> J. Sun,<sup>55</sup> J. Sziklai,<sup>61</sup> E.M. Takagui,<sup>52</sup> A. Takahara,<sup>10</sup> A. Taketani,<sup>48,49</sup> Y. Tanaka,<sup>41</sup> S. Taneja,<sup>55</sup> K. Tanida,<sup>49,53</sup> M.J. Tannenbaum,<sup>6</sup> S. Tarafdar,<sup>2</sup> A. Taranenko,<sup>54</sup> E. Tennant,<sup>43</sup> H. Themann,<sup>55</sup> T. Todoroki,<sup>48,58</sup> L. Tomášek,<sup>24</sup> M. Tomášek,<sup>13,24</sup> H. Torii,<sup>21</sup> R.S. Towell,<sup>1</sup> I. Tserruya,<sup>60</sup> Y. Tsuchimoto,<sup>10</sup> T. Tsuji,<sup>10</sup> C. Vale,<sup>6</sup> H.W. van Hecke,<sup>34</sup> M. Vargyas,<sup>16</sup> E. Vazquez-Zambrano,<sup>12</sup> A. Veicht,<sup>12</sup> J. Velkovska,<sup>59</sup> R. Vértesi,<sup>61</sup> M. Virius,<sup>13</sup> A. Vossen,<sup>23</sup> V. Vrba,<sup>13,24</sup> E. Vznuzdaev,<sup>47</sup> X.R. Wang,<sup>43</sup> D. Watanabe,<sup>21</sup> K. Watanabe,<sup>58</sup> Y. Watanabe,<sup>48,49</sup> Y.S. Watanabe,<sup>10</sup> F. Wei,<sup>25</sup> R. Wei,<sup>54</sup> S.N. White,<sup>6</sup> D. Winter,<sup>12</sup> S. Wolin,<sup>23</sup> C.L. Woody,<sup>6</sup> M. Wysocki,<sup>11</sup> Y.L. Yamaguchi,<sup>10</sup> R. Yang,<sup>23</sup> A. Yanovich,<sup>22</sup> J. Ying,<sup>19</sup> S. Yokkaichi,<sup>48,49</sup> Z. You,<sup>34</sup> I. Younus,<sup>42</sup> I.E. Yushmanov,<sup>30</sup> W.A. Zajc,<sup>12</sup> A. Zelenski,<sup>5</sup> and L. Zolin<sup>26</sup>

## (PHENIX Collaboration)

- <sup>1</sup>Abilene Christian University, Abilene, Texas 79699, USA  
<sup>2</sup>Department of Physics, Banaras Hindu University, Varanasi 221005, India  
<sup>3</sup>Bhabha Atomic Research Centre, Bombay 400 085, India  
<sup>4</sup>Baruch College, City University of New York, New York, New York, 10010 USA  
<sup>5</sup>Collider-Accelerator Department, Brookhaven National Laboratory, Upton, New York 11973-5000, USA  
<sup>6</sup>Physics Department, Brookhaven National Laboratory, Upton, New York 11973-5000, USA  
<sup>7</sup>University of California - Riverside, Riverside, California 92521, USA  
<sup>8</sup>Charles University, Ovocný trh 5, Praha 1, 116 36, Prague, Czech Republic  
<sup>9</sup>Chonbuk National University, Jeonju, 561-756, Korea  
<sup>10</sup>Center for Nuclear Study, Graduate School of Science, University of Tokyo, 7-3-1 Hongo, Bunkyo, Tokyo 113-0033, Japan  
<sup>11</sup>University of Colorado, Boulder, Colorado 80309, USA  
<sup>12</sup>Columbia University, New York, New York 10027 and Nevis Laboratories, Irvington, New York 10533, USA  
<sup>13</sup>Czech Technical University, Zikova 4, 166 36 Prague 6, Czech Republic  
<sup>14</sup>Dapnia, CEA Saclay, F-91191, Gif-sur-Yvette, France  
<sup>15</sup>Debrecen University, H-4010 Debrecen, Egyetem tér 1, Hungary  
<sup>16</sup>ELTE, Eötvös Loránd University, H - 1117 Budapest, Pázmány P. s. 1/A, Hungary  
<sup>17</sup>Ewha Womans University, Seoul 120-750, Korea  
<sup>18</sup>Florida State University, Tallahassee, Florida 32306, USA  
<sup>19</sup>Georgia State University, Atlanta, Georgia 30303, USA  
<sup>20</sup>Hanyang University, Seoul 133-792, Korea  
<sup>21</sup>Hiroshima University, Kagamiyama, Higashi-Hiroshima 739-8526, Japan  
<sup>22</sup>IHEP Protvino, State Research Center of Russian Federation, Institute for High Energy Physics, Protvino, 142281, Russia  
<sup>23</sup>University of Illinois at Urbana-Champaign, Urbana, Illinois 61801, USA  
<sup>24</sup>Institute of Physics, Academy of Sciences of the Czech Republic, Na Slovance 2, 182 21 Prague 8, Czech Republic  
<sup>25</sup>Iowa State University, Ames, Iowa 50011, USA  
<sup>26</sup>Joint Institute for Nuclear Research, 141980 Dubna, Moscow Region, Russia  
<sup>27</sup>Helsinki Institute of Physics and University of Jyväskylä, P.O.Box 35, FI-40014 Jyväskylä, Finland  
<sup>28</sup>KEK, High Energy Accelerator Research Organization, Tsukuba, Ibaraki 305-0801, Japan  
<sup>29</sup>Korea University, Seoul, 136-701, Korea  
<sup>30</sup>Russian Research Center "Kurchatov Institute", Moscow, 123098 Russia  
<sup>31</sup>Kyoto University, Kyoto 606-8502, Japan  
<sup>32</sup>Laboratoire Leprince-Ringuet, Ecole Polytechnique, CNRS-IN2P3, Route de Saclay, F-91128, Palaiseau, France  
<sup>33</sup>Lawrence Livermore National Laboratory, Livermore, California 94550, USA  
<sup>34</sup>Los Alamos National Laboratory, Los Alamos, New Mexico 87545, USA  
<sup>35</sup>LPC, Université Blaise Pascal, CNRS-IN2P3, Clermont-Fd, 63177 Aubiere Cedex, France  
<sup>36</sup>Department of Physics, Lund University, Box 118, SE-221 00 Lund, Sweden  
<sup>37</sup>University of Maryland, College Park, Maryland 20742, USA  
<sup>38</sup>Department of Physics, University of Massachusetts, Amherst, Massachusetts 01003-9337, USA  
<sup>39</sup>Muhlenberg College, Allentown, Pennsylvania 18104-5586, USA  
<sup>40</sup>Myongji University, Yongin, Kyonggido 449-728, Korea  
<sup>41</sup>Nagasaki Institute of Applied Science, Nagasaki-shi, Nagasaki 851-0193, Japan  
<sup>42</sup>University of New Mexico, Albuquerque, New Mexico 87131, USA  
<sup>43</sup>New Mexico State University, Las Cruces, New Mexico 88003, USA  
<sup>44</sup>Department of Physics and Astronomy, Ohio University, Athens, Ohio 45701, USA  
<sup>45</sup>Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831, USA  
<sup>46</sup>IPN-Orsay, Université Paris Sud, CNRS-IN2P3, BP1, F-91406, Orsay, France  
<sup>47</sup>PNPI, Petersburg Nuclear Physics Institute, Gatchina, Leningrad region, 188300, Russia  
<sup>48</sup>RIKEN Nishina Center for Accelerator-Based Science, Wako, Saitama 351-0198, Japan  
<sup>49</sup>RIKEN BNL Research Center, Brookhaven National Laboratory, Upton, New York 11973-5000, USA  
<sup>50</sup>Physics Department, Rikkyo University, 3-34-1 Nishi-Ikebukuro, Toshima, Tokyo 171-8501, Japan  
<sup>51</sup>Saint Petersburg State Polytechnic University, St. Petersburg, 195251 Russia  
<sup>52</sup>Universidade de São Paulo, Instituto de Física, Caixa Postal 66318, São Paulo CEP05315-970, Brazil  
<sup>53</sup>Department of Physics and Astronomy, Seoul National University, Seoul, Korea  
<sup>54</sup>Chemistry Department, Stony Brook University, SUNY, Stony Brook, New York 11794-3400, USA  
<sup>55</sup>Department of Physics and Astronomy, Stony Brook University, SUNY, Stony Brook, New York 11794-3400, USA  
<sup>56</sup>University of Tennessee, Knoxville, Tennessee 37996, USA  
<sup>57</sup>Department of Physics, Tokyo Institute of Technology, Oh-okayama, Meguro, Tokyo 152-8551, Japan  
<sup>58</sup>Institute of Physics, University of Tsukuba, Tsukuba, Ibaraki 305, Japan  
<sup>59</sup>Vanderbilt University, Nashville, Tennessee 37235, USA  
<sup>60</sup>Weizmann Institute, Rehovot 76100, Israel  
<sup>61</sup>Institute for Particle and Nuclear Physics, Wigner Research Centre for Physics, Hungarian Academy of Sciences (Wigner RCP, RMKI) H-1525 Budapest 114, POBox 49, Budapest, Hungary

<sup>62</sup>Yonsei University, IPAP, Seoul 120-749, Korea  
<sup>63</sup>Chonbuk

(Dated: April 9, 2012)

Neutral-pion,  $\pi^0$ , spectra were measured at midrapidity ( $|y| < 0.35$ ) in Au+Au collisions at  $\sqrt{s_{NN}} = 39$  and 62.4 GeV and compared to earlier measurements at 200 GeV in the  $1 < p_T < 10$  GeV/c transverse-momentum ( $p_T$ ) range. The high- $p_T$  tail is well described by a power law in all cases and the powers decrease significantly with decreasing center-of-mass energy. The change of powers is very similar to that observed in the corresponding  $p+p$ -collision spectra. The nuclear-modification factors ( $R_{AA}$ ) show significant suppression and a distinct energy dependence at moderate  $p_T$  in central collisions. At high  $p_T$ ,  $R_{AA}$  is similar for 62.4 and 200 GeV at all centralities. Perturbative-quantum-chromodynamics calculations that describe  $R_{AA}$  well at 200 GeV, fail to describe the 39 GeV data, raising the possibility that the relative importance of initial-state effects and soft processes increases at lower energies. A conclusion that the region where hard processes are dominant is reached only at higher  $p_T$ , is also supported by the  $x_T$  dependence of the  $x_T$ -scaling power-law exponent.

PACS numbers: 25.75.Dw

Large transverse-momentum ( $p_T$ ) particles produced in high-energy nucleus-nucleus (AB) collisions play a crucial role in studying the properties of the medium created in relativistic heavy-ion collisions. Most hadrons at sufficiently high  $p_T$  are fragmentation products of hard-scattered partons and their production rate in vacuum, as measured in  $p+p$  collisions, is well described by perturbative quantum chromodynamics (pQCD) [1]. In the absence of any nuclear effects the production rate in relativistic heavy-ion collisions in the pQCD regime, i.e. at sufficiently high  $p_T$ , would scale with the increased probability that a hard scattering occurs, due to the large number of nucleons. This probability is characterized by the nuclear thickness function  $T_{AB}$  [2]. However, such scaling has been violated to various degrees depending both on collision energy,  $\sqrt{s_{NN}}$ , and hadron  $p_T$ . At lower collision energies, the hadron yield is enhanced above the expected scaling. This was first observed in  $p+A$  and this enhancement is generally attributed to multiple soft scattering (“Cronin effect” [3]), and is presumed to occur in ion-ion collisions as well. Initial parton distribution functions in nuclei (nPDF) are different from those in protons [4].

Finally, if a dense, colored medium is formed in the AB collision, the hard-scattered parton may traverse some of it, losing energy in the process. Therefore, the observed yield at a given (high)  $p_T$  will be lower than that expected from  $T_{AB}$  scaling, exhibiting “suppression” or “jet quenching,” described in terms of the nuclear-modification factor,  $R_{AA}$  (see Eq. (1)). Alternatively, other studies divide the yields for heavy-ion collisions at one energy with those for the same colliding species at a lower energy Au+Au, rather than scaled  $p+p$  reference data, to study energy and centrality scaling [5].

One of the first discoveries at the Relativistic Heavy-Ion Collider (RHIC) was a very large hadron suppression at high  $p_T$  (above  $\approx 3$  GeV/c) in  $\sqrt{s_{NN}} = 130$  and 200 GeV Au+Au collisions [6–9]. This suppression was attributed to the dominance of parton energy loss in the medium, i.e. to final state effects. To test this hypothesis, the same

measurements were performed in  $d+Au$  collisions [10], where the formation of the hot, dense partonic medium is not expected, and initial-state effects (if any) prevail. No suppression in  $d+Au$  data was observed leaving little (if any) room for the initial-state effects as the origin of the large jet quenching observed in Au + Au. Studies with the lighter Cu+Cu system at three energies ( $\sqrt{s_{NN}} = 22.4$ , 62.4 and 200 GeV [11]) have revealed that at  $\sqrt{s_{NN}} = 22.4$  GeV mechanisms that enhance  $R_{AA}$  ( $> 1$ ) dominate at all centralities. Note, however, that this data set had very limited  $p_T$  range ( $p_T < 4$  GeV/c). At 62.4 GeV, jet quenching overwhelms any enhancement and leads to a suppression ( $R_{AA} < 1$ ) in more central collisions.

The low-energy scan at RHIC provides an opportunity to study the transition from enhancement ( $R_{AA} > 1$ ) to suppression ( $R_{AA} < 1$ ) and the evolution of  $R_{AA}$  with collision energy, centrality and  $p_T$ . The results put constraints on energy-loss models (see [12] and references therein).

Here, we present new measurements by the PHENIX experiment at RHIC of  $\pi^0$  invariant yields and  $R_{AA}$  in Au + Au collisions at  $\sqrt{s_{NN}} = 39$  and 62.4 GeV. The data were taken during the 2010 run and the  $p_T$  limits (statistics) were 8 GeV/c and 10 GeV/c, respectively. Reference  $p+p$ -collision data for  $\sqrt{s_{NN}} = 62.4$  GeV were taken in the same experiment in the 2006 run [1], while for  $\sqrt{s_{NN}} = 39$  GeV, data measured in the FERMILAB experiment E706 were used [13].

Neutral pions were measured on a statistical basis via their  $\pi^0 \rightarrow \gamma\gamma$  decay branch with the electromagnetic calorimeter (EMCal) [14]. The EMCal comprises two calorimeter types: 6 sectors of lead scintillator sampling calorimeter (PbSc) and 2 sectors of lead glass Čerenkov calorimeter (PbGl). Each sector is located  $\approx 5$  m from the beamline and subtends  $|\eta| < 0.35$  in pseudorapidity and  $\Delta\phi = 22.5^\circ$  in azimuth. This Letter presents results obtained with the PbSc sectors only. The segmentation of the PbSc ( $\Delta\eta \times \Delta\phi = 0.01 \times 0.01$ ) ensures that the two photons from the  $\pi^0 \rightarrow \gamma\gamma$  decays are very well resolved

up to  $p_T < 12 \text{ GeV}/c$ , i.e. across the entire  $p_T$  range of this measurement.

The results are based on data sets of  $3.5 \cdot 10^8$  and  $7.0 \cdot 10^8$  minimum bias Au + Au events at 39 and 62.4 GeV, respectively. The minimum bias (MB) trigger for both  $\sqrt{s_{NN}} = 39$  and 62.4 GeV was provided by the Beam-Beam-Counters (BBC) [15], located close to the beam axis in both directions and covering  $3.0 \leq |\eta| \leq 3.9$ . In order to reduce background at least two hits were required in both BBC's. This condition selects  $\sim 86\%$  of the total inelastic cross section. The centrality selection in Au + Au collisions at both energies was based on the charged signal sum of the BBC's, which is proportional to the charged particle multiplicity. For each centrality the average number of binary collisions ( $\langle N_{\text{coll}} \rangle$ ) and the number of participants ( $\langle N_{\text{part}} \rangle$ ) were calculated using a Glauber model [2] based Monte Carlo code.

TABLE I: Sources of systematic uncertainties and their relative effect (in %) on the invariant yields

$p_T$ energy	2 GeV/c      5 GeV/c				Type
	39 GeV (62.4 GeV)				
Yield extraction	3%	(3%)	3%	(3%)	A
PID efficiency	4.5%	(4.5%)	4.5%	(4.5%)	B
Energy scale	10.5%	(8.0%)	14.5%	(10.0%)	B
Acceptance	2%	(2%)	2%	(2%)	B
Conversion	4%	(4%)	4%	(4%)	B
Off vertex	1.5%	(1.5%)	1.5%	(1.5%)	C
Total for $\pi^0$ yields	12.7%	(10.7%)	16.2%	(12.3%)	

The PHENIX analysis of neutral pions is described in detail elsewhere [9]. Table I lists the sources of systematic uncertainties on the extracted- $\pi^0$  invariant yields in this analysis. They can be divided into three different categories: (1) Type-A,  $p_T$ -uncorrelated; (2) Type-B,  $p_T$ -correlated, where the correlation may be an arbitrary smooth function; (3) Type-C,  $p_T$ -correlated, where all points move by the same fraction up or down. The main sources of systematic uncertainties in the  $\pi^0$  measurement are the energy scale, yield extraction and particle-identification (PID) efficiency correction.

Figure 1 shows the invariant yields of the  $\pi^0$ s for all centralities and also in minimum bias collisions. From fitting the  $\sqrt{s_{NN}} = 39$  and 62.4 GeV minimum bias spectra with a power law function ( $\propto p_T^n$ ) for  $p_T > 4 \text{ GeV}/c$ , we obtained powers  $n_{39} = -13.04 \pm 0.08$  and  $n_{62.4} = -10.60 \pm 0.03$ , respectively, significantly steeper than at  $\sqrt{s_{NN}} = 200 \text{ GeV}$ , where  $n_{200} = -8.06 \pm 0.01$  for MB collisions [9]. The slopes of the corresponding  $p+p$ -collision spectra are somewhat different, but comparable,  $n_{39}^{pp} = -13.59 \pm 0.21$ ,  $n_{62.4}^{pp} = -9.82 \pm 0.18$  and  $n_{200}^{pp} = -8.22 \pm 0.09$ , respectively.

Nuclear effects on the  $\pi^0$  production are quantified us-

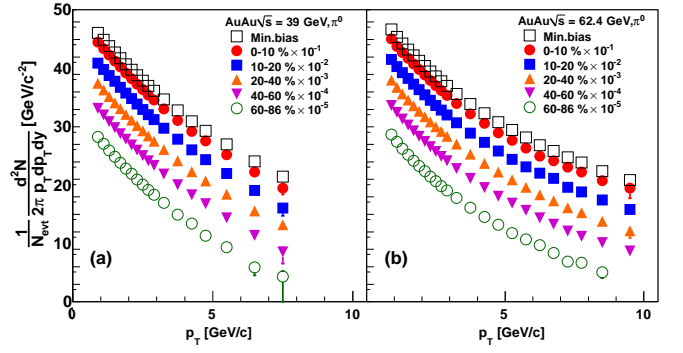


FIG. 1: (Color online) Invariant yields of  $\pi^0$  in Au + Au at  $\sqrt{s_{NN}} = 39 \text{ GeV}$  (a) and  $62.4 \text{ GeV}$  (b) in all centralities and minimum bias. Only statistical uncertainties are shown.

ing the nuclear modification factor

$$R_{AA}(p_T) = \frac{(1/N_{AA}^{\text{evt}}) d^2 N_{AA}^{\pi^0} / dp_T dy}{\langle T_{AB} \rangle \times d^2 \sigma_{pp}^{\pi^0} / dp_T dy}, \quad (1)$$

where  $\sigma_{pp}^{\pi^0}$  is the production cross section of  $\pi^0$  in  $p+p$  collisions, and  $\langle T_{AB} \rangle = \langle N_{\text{coll}} \rangle / \sigma_{pp}^{\text{inel}}$  is the nuclear thickness function averaged over the range of impact parameters contributing to the given centrality class according to the Glauber model. Thus  $R_{AA}$  compares the yield observed in  $A + A$  collisions to the yield expected from the superposition of  $N_{\text{coll}}$  independent  $p+p$  interactions. In the absence of nuclear effects,  $R_{AA}$  should be equal to unity. However,  $R_{AA} \approx 1$  does not necessarily imply the absence of suppression, it may also indicate a balance between enhancing and depleting mechanisms.

In order to calculate  $R_{AA}$ , a reference  $p_T$  distribution in  $p+p$  collisions is needed. Preferably this is measured with the same detector, in which case many systematic uncertainties cancel in the ratio. The PHENIX experiment has measured the  $\pi^0$  cross section in  $p+p$  collisions at  $\sqrt{s_{NN}} = 62.4 \text{ GeV}$  [1] but only up to  $p_T = 7 \text{ GeV}/c$  while the current Au+Au measurement reaches up to  $10 \text{ GeV}/c$ . Hence the  $p+p$  data were fitted with a power law function between  $4.5 < p_T < 7 \text{ GeV}/c$  and then extrapolated. The systematic uncertainty resulting from this extrapolation reaches 20% at  $10 \text{ GeV}/c$ , estimated from a series of fits, where each time one or more randomly selected points are omitted and the remaining points are re-fitted.

So far PHENIX has not measured the  $p+p$  spectrum of  $\pi^0$  at  $\sqrt{s_{NN}} = 39 \text{ GeV}$ . Therefore, data from the Fermilab experiment E706 [13] were used. However, the E706 acceptance ( $-1. < |\eta| < 0.5$ ) is different from that of PHENIX ( $|\eta| < 0.35$ ), and since  $dN/d\eta$  is not flat, a  $p_T$ -dependent correction was applied to the E706 data. This correction factor was determined from a PYTHIA simulation by means of the ratio of yields (normalized per unit rapidity) when calculated from the observed yield in the

PHENIX and E706 acceptance windows. The systematic uncertainty of the correction is 1–2% at 3 GeV/c but reaches 20% at 8 GeV/c.

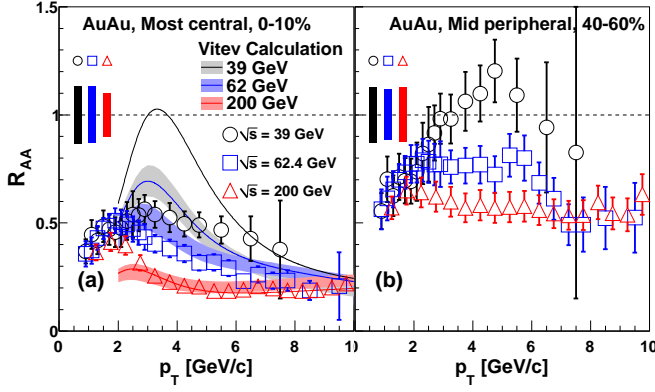


FIG. 2: (Color online) Nuclear modification factor ( $R_{AA}$ ) of  $\pi^0$  in Au + Au collisions in most central 0–10% (a) and mid-peripheral 40–60% (b). Error bars are the quadratic sum of statistical and  $p_T$ -correlated systematic uncertainties (including systematic uncertainties from the  $p+p$ -collision reference). Boxes around 1 are the quadratic sum of the C-type uncertainties combined with the  $N_{coll}$  uncertainties. These are fully correlated between different energies. Also shown for central collisions are pQCD calculations [16] with regular Cronin-effect (solid lines) and with the Cronin-effect reduced by a factor of two for all three energies (bands).

Figure 2 shows the nuclear modification factor of  $\pi^0$ s measured in Au + Au collisions at  $\sqrt{s_{NN}} = 39, 62.4$  and 200 GeV (data from [9]) as a function of  $p_T$  for most central collisions (a) and 40–60% centrality (b). In the most central collisions (0–10%) there is a significant suppression for all three energies, while in mid-peripheral collisions (40–60%) at  $\sqrt{s_{NN}} = 39$  GeV,  $R_{AA}$  is consistent with unity above  $p_T > 3$  GeV/c.

For 0–10% pQCD calculations [16, 17] are also shown. The solid curves are calculated with the same parametrization that was successful for 200 GeV Au + Au data (and also 200 GeV Cu + Cu [11]). Neither the 62.4, nor the 39 GeV data are consistent with the predictions. The only qualitative agreement is that the turnover point of the  $R_{AA}$  curves moves to higher  $p_T$  with lower collision energy as observed in the data. The bands are calculated within the same framework but with the Cronin-effect reduced and the energy loss varied by  $\pm 10\%$ . The 200 GeV data are still well described, the 62.4 GeV data are consistent within uncertainties, but the 39 GeV  $R_{AA}$ , particularly the shape, is inconsistent with the corresponding band.

Coupled with the observations that the slopes at high  $p_T$  become much steeper, but the bulk properties (like elliptic flow, energy density, apparent temperature) change only slowly in the collision energy range in question, it is quite conceivable that hard scattering as a source of par-

ticles at a given  $p_T$  becomes completely dominant only at higher transverse momentum, i.e. jet quenching will be “masked” up to higher  $p_T$ . Note that while the shapes at lower  $p_T$  are different, at  $p_T \gtrsim 7$  GeV/c  $R_{AA}$  is essentially the same for the 62.4 and 200 GeV data, irrespective of centrality (see also Fig. 3). While in the 39 GeV data  $R_{AA}$  also shows a decreasing trend at higher  $p_T$ , unfortunately the  $p_T$  reach of the current data sample precludes any conclusion as to what would happen to their  $R_{AA}$  at even higher  $p_T$ .

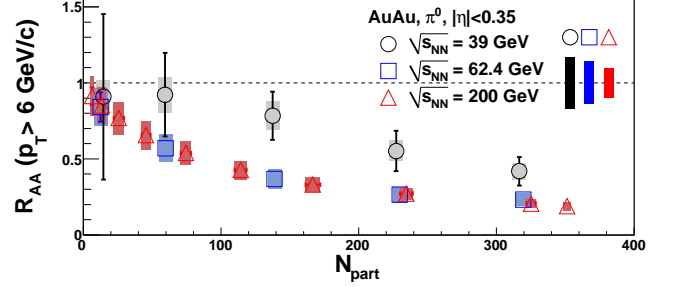


FIG. 3: (Color online) Nuclear modification factor averaged for  $p_T > 6$  GeV/c. Uncertainties are shown as error bars (statistical), boxes (sum of  $p_T$ -uncorrelated and  $N_{coll}$ ), boxes around one (Type B and C and uncertainties from the  $p+p$ -collision reference).

Figure 3 shows  $p_T$ -averaged  $R_{AA}$  as a function of the number of participants. The averaging was done above  $p_T > 6$  GeV/c. Our first observation is that  $R_{AA}$  decreases with increasing centrality even for the lowest-energy system. Similarly, as already discussed in the context of Fig. 2, at high enough  $p_T$  the suppression is the same at 62.4 and 200 GeV, at all centralities. This is remarkable because the power  $n$  of the fit to the spectra changes approximately by two units from 200 to 62.4 GeV, so the average momentum loss of the partons also has to be different in order to compensate the effect of the changing slope. The average momentum loss is usually defined by the fractional momentum shift  $\delta p_T/p_T$  between the corresponding Au + Au and  $T_{AA}$ -scaled  $p+p$  spectra as follows. Since the power law tails of the  $p+p$  and Au + Au spectra are similar, they can be fitted simultaneously with the same function and same power  $n$

$$f(p_T) = \frac{A}{(p_T(1 + \delta p_T/p_T))^n} \quad (2)$$

with  $\delta p_T$  being the horizontal shift between the scaled  $p+p$  and the Au + Au spectra. In panel (a) of Fig. 4, the observed fractional momentum shifts are shown for central collisions, as a function of the Au + Au  $p_T$ .

Inclusive single-particle spectra at sufficiently high  $p_T$  and collision energy were predicted to exhibit scaling with the variable  $x_T = 2p_T/\sqrt{s}$  such that the produc-

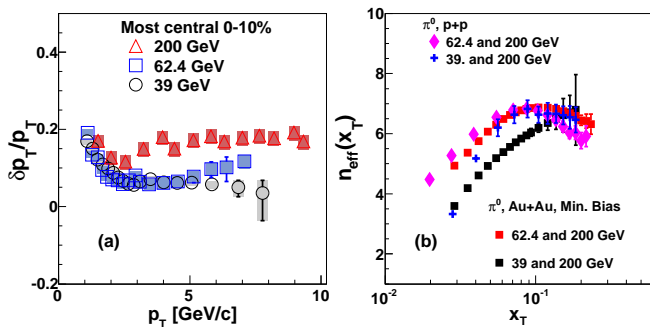


FIG. 4: (Color online) (a) Fractional momentum shift  $\delta p_T / p_T$  between Au + Au and  $T_{AA}$ -scaled  $p+p$  data as a function of the Au + Au  $p_T$ . (b) Power  $n_{\text{eff}}$  of  $x_T$ -scaling for  $p+p$  and Au + Au (minimum bias) at various collision energies.

tion cross section can be written in a form [18, 19]

$$E \frac{d^3\sigma}{dp^3} = \frac{1}{\sqrt{s}^{n(x_T, \sqrt{s})}} G(x_T) \quad (3)$$

where  $G(x_T)$  is a universal function and  $n(x_T, \sqrt{s})$  characterizes the specific process [19]. The scaling power  $n_{\text{eff}}(x_T)$  between any pair of  $\sqrt{s_{NN}}$  energies is then calculated as

$$n_{\text{eff}}(x_T) = \frac{\log(\text{Yield}(x_T, \sqrt{s_1}) / \text{Yield}(x_T, \sqrt{s_2}))}{\log(\sqrt{s_2} / \sqrt{s_1})} \quad (4)$$

In panel (b) of Fig. 4,  $n_{\text{eff}}(x_T)$  is shown when comparing invariant- $\pi^0$  yields in  $p+p$  and Au + Au collisions at different energies. Both the shape and the magnitude of  $n_{\text{eff}}(x_T)$  is similar for the 62.4/200 GeV  $p+p$  and Au + Au as well as for the 39/200 GeV  $p+p$  data. The rise of  $n_{\text{eff}}(x_T)$  at lower  $x_T$  can be attributed to the dominance of soft processes [20], while at higher  $x_T$  they deviate strongly from leading-twist scaling predictions [19, 21]. However, the shape of  $n_{\text{eff}}(x_T)$  in the 39 and 200 GeV Au + Au comparison is very different from all others. It may not even reach its maximum in the measured  $x_T$  range, and its constant rise is similar to the rise observed in the low- $x_T$  (soft) region of the other data shown. One possible explanation could be that while present, hard scattering is still not the overwhelming source of high- $p_T$   $\pi^0$ s in the currently available  $p_T$  range in 39 GeV Au + Au collisions.

In summary, the  $\pi^0$   $p_T$  spectra were measured in Au + Au collisions at two different energies,  $\sqrt{s_{NN}} = 39$  and 62.4 GeV, and compared to the earlier result for  $\sqrt{s_{NN}} = 200$  GeV. In all cases the high  $p_T$  part of the invariant yields can be well described with a single power law function. The powers decrease considerably at lower  $\sqrt{s_{NN}}$ , and since the soft processes change only slowly with collision energy, jet quenching might be “masked” up to higher transverse momenta. The high- $p_T$   $\pi^0$  yields in Au + Au at 62.4 GeV are suppressed,

and above  $p_T > 6$  GeV/ $c$  the data points are comparable with the 200 GeV results at all centralities. The  $\pi^0$  yields in Au + Au at 39 GeV are suppressed in the most central collisions, but no suppression is apparent in more peripheral collisions. At lower energies, a decreasing momentum shift compensates for the steeper slopes at high  $p_T$ , making the  $R_{AA}$ ’s comparable, in fact, identical in the case of 62.4 and 200 GeV. When related to 200 GeV,  $n_{\text{eff}}(x_T)$  is similar for 62.4 and 39 GeV  $p+p$  and 62.4 GeV Au + Au, but very different for the 39 GeV Au + Au data. The new data provided in a wide energy-range of Au + Au collisions will help to constrain energy-loss models.

We thank the staff of the Collider-Accelerator and Physics Departments at Brookhaven National Laboratory and the staff of the other PHENIX participating institutions for their vital contributions. We acknowledge support from the Office of Nuclear Physics in the Office of Science of the Department of Energy, the National Science Foundation, Abilene Christian University Research Council, Research Foundation of SUNY, and Dean of the College of Arts and Sciences, Vanderbilt University (U.S.A), Ministry of Education, Culture, Sports, Science, and Technology and the Japan Society for the Promotion of Science (Japan), Conselho Nacional de Desenvolvimento Científico e Tecnológico and Fundação de Amparo à Pesquisa do Estado de São Paulo (Brazil), Natural Science Foundation of China (P. R. China), Ministry of Education, Youth and Sports (Czech Republic), Jyväskylä University (Finland), Centre National de la Recherche Scientifique, Commissariat à l’Énergie Atomique, and Institut National de Physique Nucléaire et de Physique des Particules (France), Ministry of Industry, Science and Technologies, Bundesministerium für Bildung und Forschung, Deutscher Akademischer Austausch Dienst, and Alexander von Humboldt Stiftung (Germany), Hungarian National Science Fund, OTKA (Hungary), Department of Atomic Energy and Department of Science and Technology (India), Israel Science Foundation (Israel), National Research Foundation and WCU program of the Ministry Education Science and Technology (Korea), Ministry of Education and Science, Russian Academy of Sciences, Federal Agency of Atomic Energy (Russia), VR and Wallenberg Foundation (Sweden), the U.S. Civilian Research and Development Foundation for the Independent States of the Former Soviet Union, the Hungarian American Enterprise Scholarship Fund, and the US-Israel Binational Science Foundation.

\* Deceased

† PHENIX Spokesperson: jacak@skipper.physics.sunysb.edu

- [1] A. Adare et al. (PHENIX Collaboration), Phys. Rev. D **79**, 012003 (2009).
- [2] M. L. Miller, K. Reygers, S. J. Sanders, and P. Steinberg,



- Annu. Rev. Nucl. Part. Sci **57**, 205 (2007).
- [3] J. W. Cronin, H. J. Frisch, and M. J. Shochet, Phys. Rev. D **11**, 3105 (1975).
  - [4] K. Eskola, H. Paukkunen, and C. Salgado, Nucl. Phys. A **855**, 150 (2011), 1011.6534.
  - [5] B. Alver et al. (PHOBOS Collaboration), Phys. Rev. C **80**, 011901 (2009).
  - [6] K. Adcox et al. (PHENIX Collaboration), Phys. Rev. Lett. **88**, 022301 (2002).
  - [7] C. Adler et al. (STAR Collaboration), Phys.Rev.Lett. **89**, 202301 (2002).
  - [8] S. S. Adler et al. (PHENIX Collaboration), Phys. Rev. Lett. **91**, 072301 (2003).
  - [9] A. Adare et al. (PHENIX Collaboration), Phys. Rev. Lett. **101**, 232301 (2008).
  - [10] S. S. Adler et al. (PHENIX Collaboration), Phys. Rev. Lett. **91**, 072303 (2003).
  - [11] A. Adare et al. (PHENIX Collaboration), Phys. Rev. Lett. **101**, 162301 (2008).
  - [12] N. Armesto et al. (2011), 1106.1106.
  - [13] L. Apanasevich et al. (E706 Collaboration), Phys. Rev. D **68**, 052001 (2003).
  - [14] L. Aphecetche et al. (PHENIX Collaboration), Nucl. Instrum. Methods A **499**, 521 (2003).
  - [15] M. Allen et al. (PHENIX Collaboration), Nucl. Instrum. Methods A **499**, 549 (2003).
  - [16] I. Vitev, private communication (2012).
  - [17] R. Sharma, I. Vitev, and B.-W. Zhang, Phys. Rev. C **80**, 054902 (2009).
  - [18] R. Blankenbecler, S. J. Brodsky, and J. F. Gunion, Phys. Lett. B **42**, 461 (1972).
  - [19] S. J. Brodsky, H. J. Pirner, and J. Raufeisenm, Phys. Lett. B **637**, 58 (2006), 0510315.
  - [20] S. M. Berman, J. D. Bjorken, and J. B. Kogut, Phys. Rev. D **4**, 3388 (1971).
  - [21] F. Arleo, S. J. Brodsky, D. S. Hwang, and A. M. Sickles, Phys. Rev. Lett. **105**, 062002 (2010).